## The geometry of spinors in Minkowski space

Daniel V. Mathews

Monash University Daniel.Mathews@monash.edu

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• attempts to explain the picture on the title slide!,



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- builds on work of Roger Penrose and Wolfgang Rindler from the 1980s on <u>Spinors and Spacetime</u>,



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Paper on arXiv soon. Email me if you want an advance copy!

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## Penrose-Rindler



General ideology: don't use vectors, use spinors for everything!

# Penrose-Rindler



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Cast of characters:

- Spinor or spin vectors: elements of C<sup>2</sup>.
- <u>Hermitian matrices</u>:  $A = A^*$ .
- Minkowski space  $\mathbb{R}^{3,1}$ : coordinates (T, X, Y, Z), metric  $dT^2 dX^2 dY^2 dZ^2$ .



$$\begin{array}{cccc} \text{Spinors} & \stackrel{\phi_1}{\longrightarrow} & \stackrel{2 \times 2 \text{ Hermitian}}{\text{matrices}} & \stackrel{\phi_2}{\longrightarrow} & \text{Minkowski space} \\ \mathbb{C}^2 & & \mathcal{H} & & \mathbb{R}^{3,1} \\ & & & & \\ & & & & \\ & & & & \\ & & & & & \\ & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & & & & \\$$

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$$\phi_1(\kappa) = \phi_1(\kappa') \Leftrightarrow \kappa = \boldsymbol{e}^{i\theta}\kappa'$$

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• Image  $\phi_1$  = Herm. matrices with det 0 & trace  $\geq$  0

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- Linear isomorphism: " $\mathcal{H} \text{ is } \mathbb{R}^{3,1}$ ", "det = norm".
- Image  $\phi (= \phi_2 \circ \phi_1) = \text{pos. light cone } L^+ (\tau^2 x^2 y^2 z^2 = 0, \tau \ge 0)$

## We understand some of the picture now



From  $\kappa \in \mathbb{C}^2$ , get a point  $\phi(\kappa) = w$  on  $L^+$ .







## Definition (Penrose-Rindler)

A <u>pointed null flag</u> is a point  $p \in L^+$  together with a 2-plane V tangent to  $L^+$  containing  $\mathbb{R}p$ .

## We understand half the picture now



From  $\kappa \in \mathbb{C}^2$ , get a point on  $L^+$  and a pointed null flag there.



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Spinoriality:

- Take  $\kappa \in \mathbb{C}^2$  and consider rotating it:  $e^{i\theta}\kappa$ .
- $\phi(e^{i\theta}\kappa)$  is constant but  $\Phi(e^{i\theta}\kappa)$  is not: plane V rotates.
- As  $\kappa$  rotates by  $\theta$ , V rotates by  $2\theta$ .

# Why you should read papers in alphabetical order



**Robert Penner** 

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• To  $w \in L^+$  associate the plane  $\langle w, x \rangle = 1$ 

• The plane intersects <u>hyperbolic space</u>  $\mathbb{H}^3$  $(\tau^2 - x^2 - y^2 - z^2 = 1, \tau > 0)$  in a <u>horosphere</u>



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Pointed null flags  $\kappa \in \mathbb{C}^2$  Penrose–Rindler Horospheres with ...  $(= p \in L^+$  and flag)

**Robert Penner** 

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There is a natural (SL(2,  $\mathbb{C}$ )-equivariant) bijection  $\mathbb{C}^2 \setminus \{0\} \longrightarrow \{\text{Horospheres in } \mathbb{H}^3 \text{ with spin directions}\}.$ 

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$$\{\kappa,\omega\} = e^{\frac{d+i\theta}{2}}$$

# New developments

#### Consequences:

- Ptolemy theorem for hyperbolic ideal tetrahedra
- New methods to compute hyperbolic structures on 3-manifolds (joint with J. Purcell)
- Equivalence of cluster algebras (Grassmannians / hyp. surfaces)
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# Thanks for listening!

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